

New 3D coded aperture collimator for X/gamma-ray wide-field imaging

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Talk #07-193

New 3D coded aperture collimator for X/gamma-ray wide-field imaging

ANIMMA 2023, Lucca, Italy

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list Contents



- 1. Introduction to X/gamma-ray imaging
- 2. X/gamma-ray imaging with coded masks
- 3. First 3D coded mask prototype
- 4. Conclusions and perspectives



Fragment of semicolumn Ionic capital. Rome, Palatine Hill, lower peristyle of Domus Augustana.

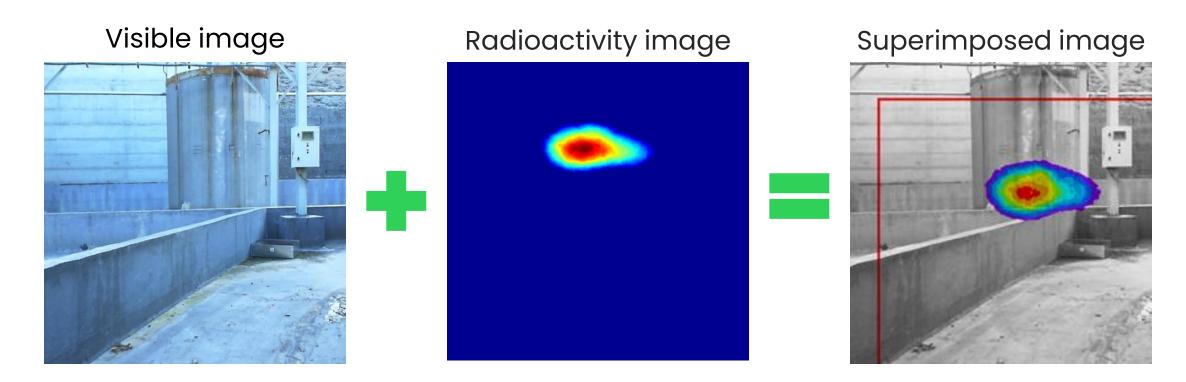
Introduction • to X/gamma-ray imaging



list Objective of X/gamma-ray imaging



Visualization of the radioactivity



dose rate, intensity, position in 2D



Applications of X/gamma-ray imaging



High-energy astrophysics

IBIS/ISGRI Simulation: Galactic Center Region 40-80 keV +10 -5 -5 -10 Galactic longitude

Credit: CEA Irfu

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Dismantling activities



Credit: S. Le-Couster/CEA

Operational radiation protection



Credit: Regis Duvignau/Reuters

Nuclear security



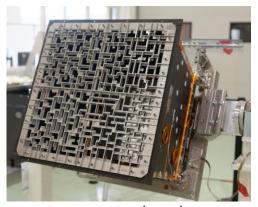
Credit: Nuclear Engineering International





Differences between applications





Credit: CNES/APC/CEA



Credit: CEA List

Parameter

Energy range

Camera size

Structure of the signal

Signal to noise ratio

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Astrophysical applications

4 keV-20 MeV

>9 m, >100 kg

Stable or transient

Weak signal, strong background

Nuclear applications

60 keV-1.3 MeV

<15 cm, <2 kg

Usually stable

Strong signal, strong background

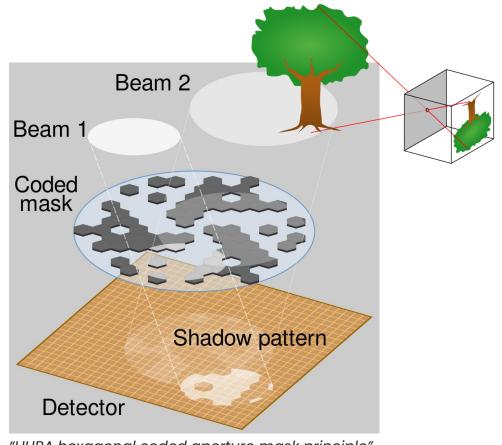
We address the two investigated applications

2 X/gamma-ray imaging with coded masks



list | Coded mask imaging concept



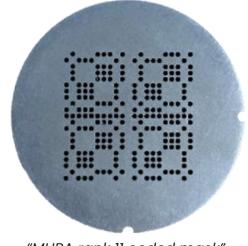


"HURA hexagonal coded aperture mask principle", by Cmglee, licensed under CC BY-SA 3.0



"Detail of the flight model of the ECLAIRs instrument's coded mask". Credit: SVOM@APC

MURA coded masks



Random array

coded masks

"MURA rank 11 coded mask". Credit: CEA List

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Objective of our work











Credit: Mirion Technologies, Inc.

Credit: NUVIA media library

Credit: CEA List

Credit: 3D PLUS

Parameter	iPIX	NuVISION	Nanopix	SpidX
Camera size	18.9 cm × 9 cm × 9 cm	10 cm × 10 cm × 24 cm	7 cm × 3 cm × 5 cm	32 cm × 16.6 cm × 10.9 cm
Coded mask dimension	1.5 cm	1.5 cm	1.5 cm	4.5 cm
Detection area	1.4 cm²	4.0 cm ²	1.4 cm²	1.4 cm ²
Field of view	48.8° × 48.8°	45° × 45°	48.8° × 48.8°	55° × 55°
Angular resolution	2.5° (150 arcmin)	3.5° (210 arcmin)	2.5° (150 arcmin)	1º (60 arcmin)

New generation of 3D coded masks

180° × 180° field of view, sub-degree angular resolution



First 3D coded mask prototype



1st 3D coded mask prototype



Parameters

Geometry

Coded pattern

Dimensions

Thickness (hole width)

Open fraction

First 3D coded mask prototype

Hemisphere

One-arm Archimedean spiral

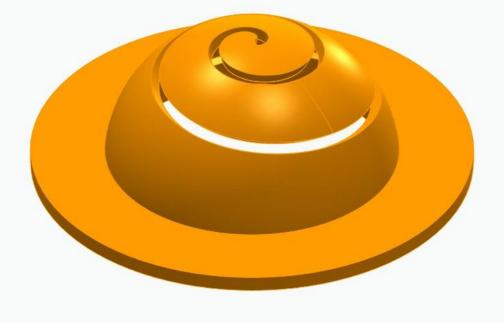
ø18.75 mm (ø28 mm with shielding)

0.75 mm

10.2 %

Stainless steel 3D printing





Tantalum machining

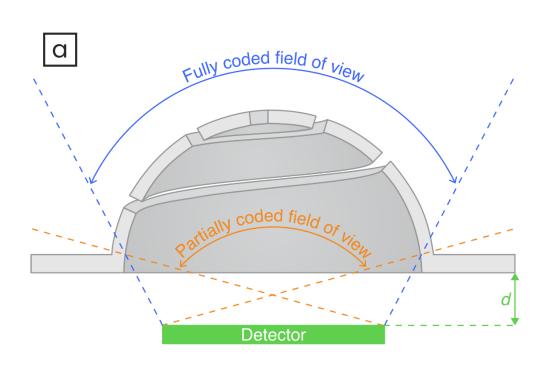


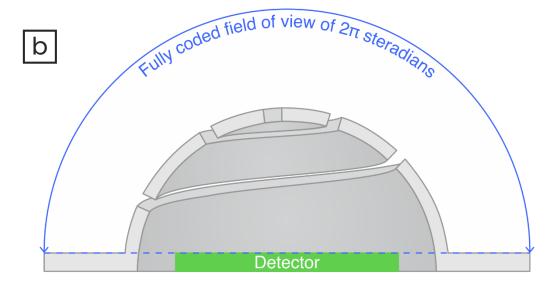


Configuration of 3D coded mask imager



Fully coded field of view and partially coded field of view of 3D coded mask imager





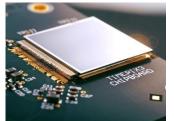
- a. Real configuration: there is a distance $d\neq 0$ between the mask and the detector
- b. Ideal configuration: mask is placed on the top of the detector d=0



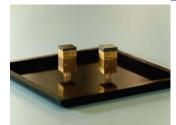
list | Our 3D coded mask imagers











Credit: CERN

Credit: CEA Irfu

Parameters	Timepix3	Caliste-HD
Dimensions	14 mm × 14 mm × 1 mm	10 mm × 10 mm × 1 mm
Pixel resolution	256 × 256	16 × 16
Pixel dimensions	625 μm × 625 μm	55 µm × 55 µm







Parameter		
Geometrical field of view		
Angular resolution		
Mask-to-detector gap		

Timepix3 imager 154.3° × 154.3° 3.5° 3.56 mm

Caliste-HD imager 126.9° × 126.9° 2.8°

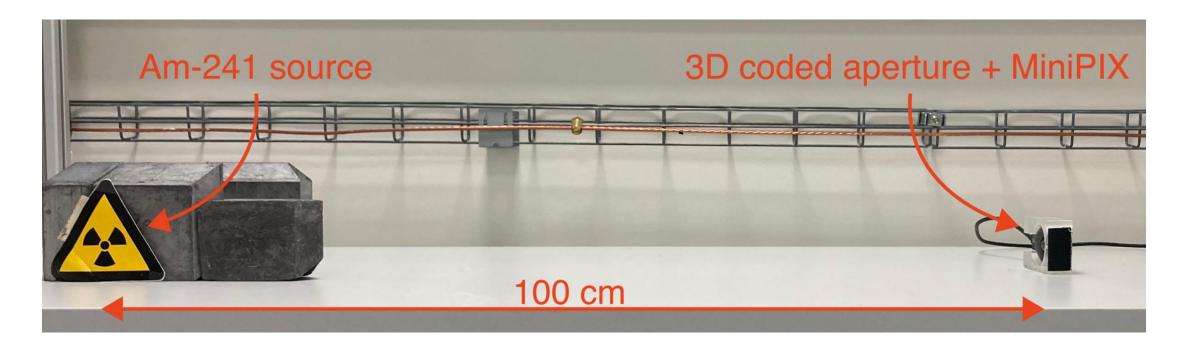
6.8 mm



Measurement with Timepix3 imager Setup



Objective: localization of an Am-241 source placed in front of the Timepix3 imager



- Am-241 activity: 74 MBq
- Acquisition time: 50 minutes
- Source-to-detector distance: 1 m



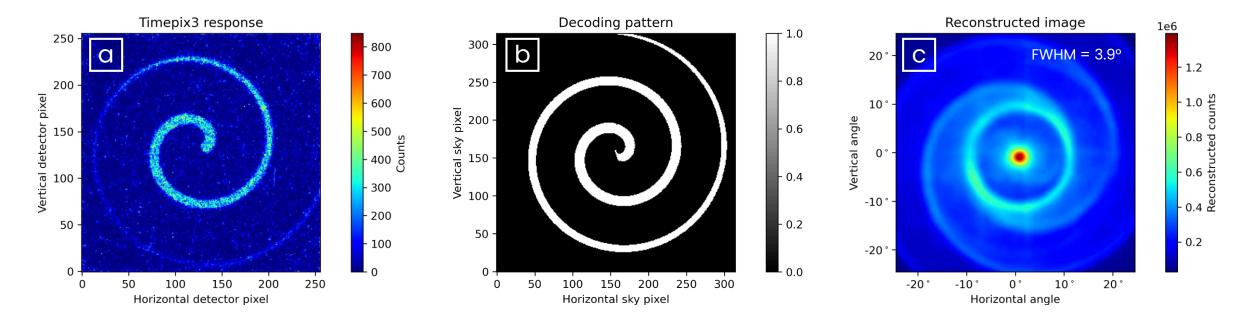


list Result

Measurement with Timepix3 imager



Am-241 (74 MBq) is localized at sky zenith



- Raw image from Timepix3
- Decoding matrix corresponding to the position of the source
- Reconstructed image of the source obtained by the deconvolution algorithm

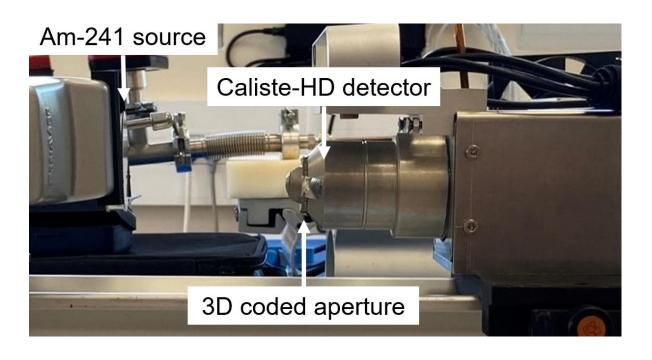
^{*} The sky is discretized with 315 \times 315 positions, with a resolution of 0.2°. Counts upper threshold = 1e+6.

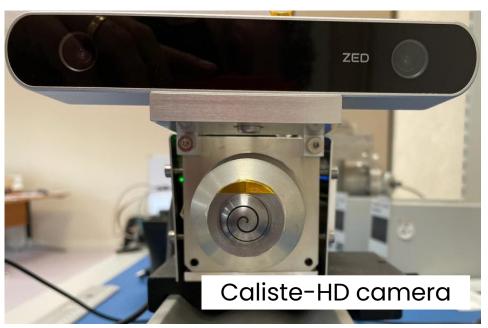


Measurement with Caliste-HD imager Setup



Objective: localization of an Am-241 source placed in front of the Caliste-HD imager





- Am-241 activity: 400 kBq
- Acquisition time: 1 hour
- Source-to-detector distance: 1 m

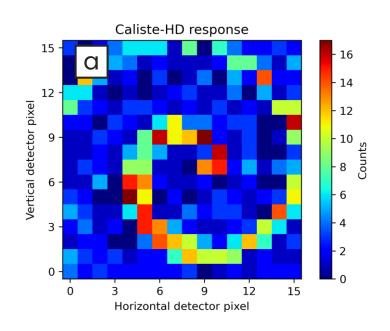


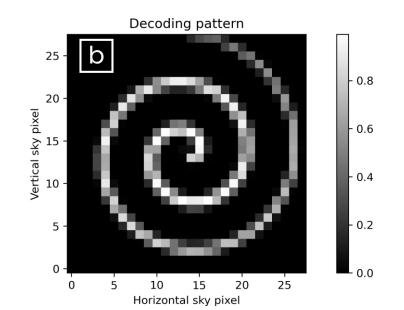


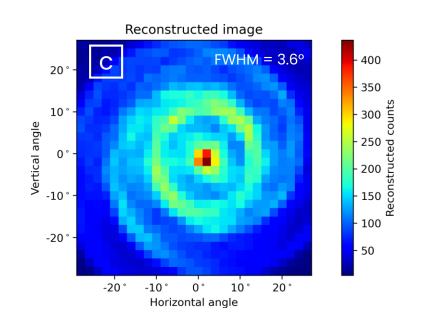
Measurement with Caliste-HD imager Result



Am-241 (400 kBq) is localized at sky zenith







- a. Raw image from Caliste-HD
- b. Decoding matrix corresponding to the position of the source
- c. Reconstructed image of the source obtained by the deconvolution algorithm



^{*} The sky is discretized with 28 × 28 positions, with a resolution of 1.8°



Improved decoding algorithms



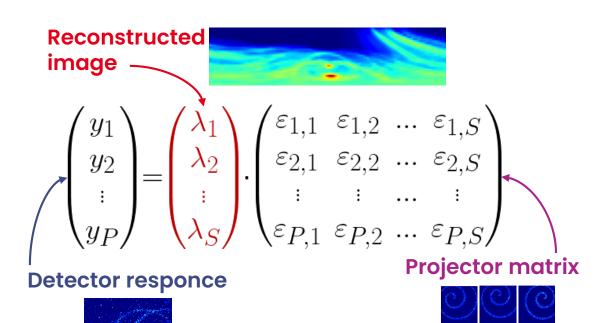
Source oriented deconvolution

Performs deconvolution and reconstructs the source with respect to its position in the sky

Decoding patterns Reconstructed **Detector** images responces

Maximum Likelihood Expectation Maximization

Iteratively estimates the Poisson data to find the optimal reconstructed image

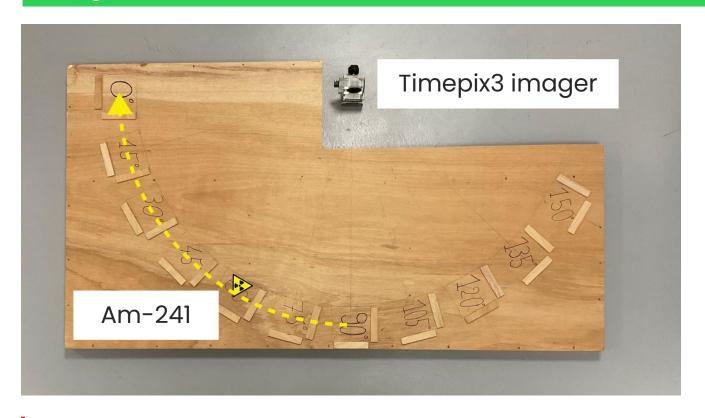


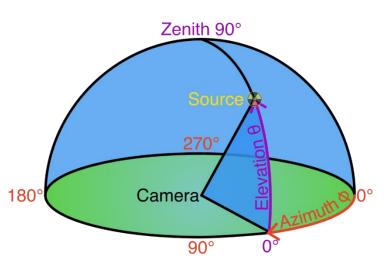


list Measurement with Timepix3 imager __ Setup



Objective: localization of an Am-241 source off-centered from the sky zenith





Horizontal coordinate system

- Am-241 activity: 74 MBq
- Acquisition time: 100 s (10 frames per second)
- Source-to-detector distance: 80 cm

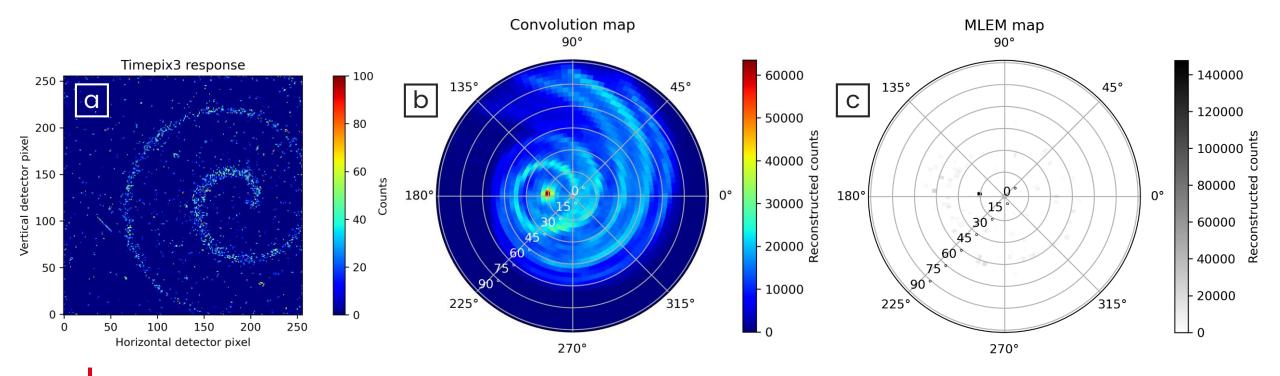




Measurement with Timepix3 imager Results



Am-241 (74 MBq) is localized at an elevation of 15°



- Raw image from Timepix3
- Reconstruction map obtained by the source oriented deconvolution
- Reconstruction map obtained by MLEM (100 iterations)



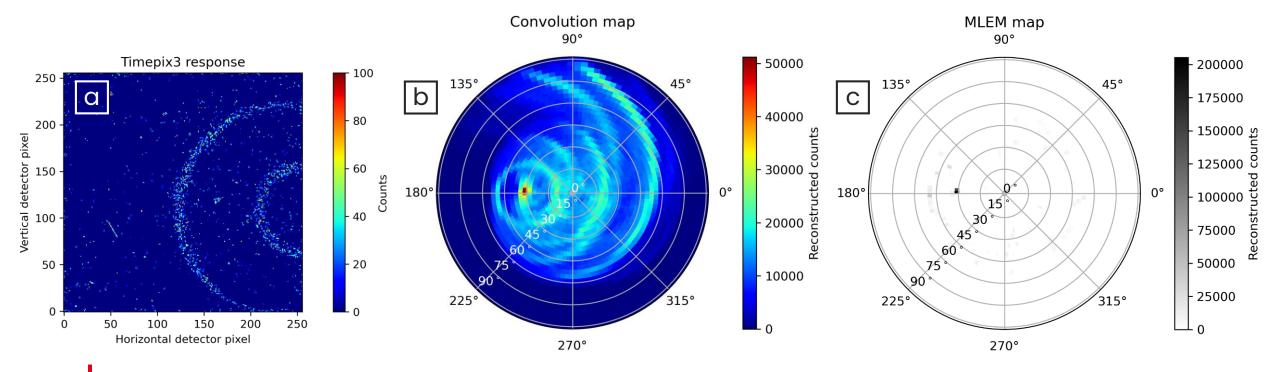
^{*} The sky is discretized with 120×30 positions, with a resolution of 3° . Counts upper threshold = 1e+2.



Measurement with Timepix3 imager Results



Am-241 (74 MBq) is localized at an elevation of 30°



- a. Raw image from Timepix3
- b. Reconstruction map obtained by the source oriented deconvolution
- c. Reconstruction map obtained by MLEM (100 iterations)



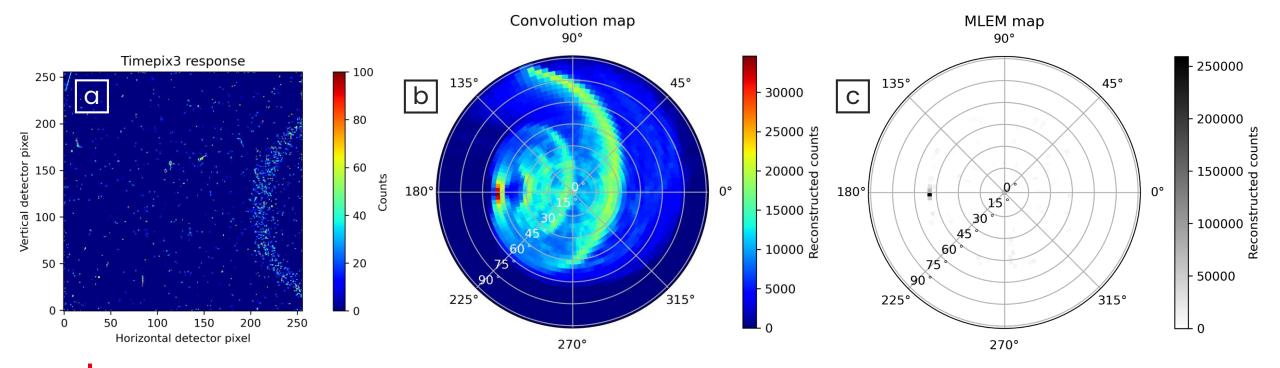
^{*} The sky is discretized with 120×30 positions, with a resolution of 3° . Counts upper threshold = 1e+2.



Measurement with Timepix3 imager Results



Am-241 (74 MBq) is localized at an elevation of 48°



- a. Raw image from Timepix3
- b. Reconstruction map obtained by the source oriented deconvolution
- c. Reconstruction map obtained by MLEM (100 iterations)



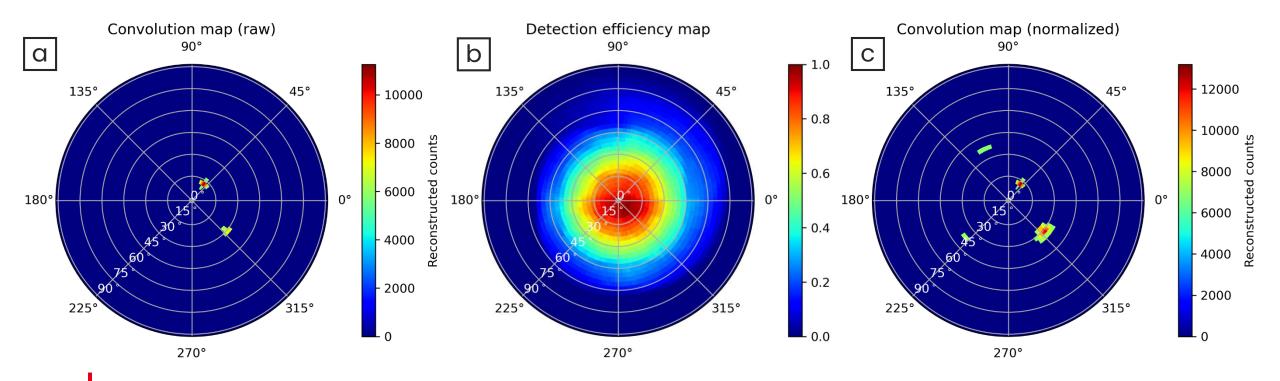
^{*} The sky is discretized with 120×30 positions, with a resolution of 3° . Counts upper threshold = 1e+2.



Simulation of two sources le+4 simulated photons for a source



Two sources are localized with equal reconstructed intensities



- a. Raw reconstruction map obtained by the source oriented deconvolution
- b. Detection efficiency map
- c. Normalized reconstruction map



^{*} The sky is discretized with 120 × 30 positions, with a resolution of 3°. Reconstructed counts lower threshold = 50%.

Conclusions and perspectives



list Conclusions



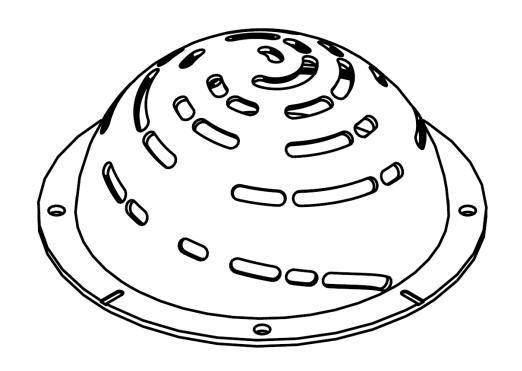
- ✓ The first 3D coded mask prototype is manufactured and charachterised
- Prototype provides larger field of view than usual commercial solutions
- ✓ Standart decoding algorithms are improved to work with complex mask geometries
- ✓ The second 3D coded mask prototype is designed and has been recently manufactured



list | Second 3D coded mask prototype



- Hemispherical geometry
- Larger mask dimensions resulting in 2π sr field of view
- Two-arm spiral pattern resulting in better sensitivity
- **Linear MURA pattern**
- "Mask/anti-mask" feature







Thank you for your attention!

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