

Physical challenges & numerical issues in controlled fusion plasmas

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► To cite this version:

Yanick Sarazin. Physical challenges & numerical issues in controlled fusion plasmas. SMAI 2021 - 10 ème Biennale Française des Mathématiques Appliquées et Industrielles, Société de Mathématiques Appliquées et Industrielles (SMAI); Institut de Mathématiques de Toulouse, Jun 2021, La Grande Motte, France. cea-03467549

HAL Id: cea-03467549 https://cea.hal.science/cea-03467549

Submitted on 6 Dec 2021

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Physical challenges & Numerical issues in Controlled fusion plasmas



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SMAI 23/06/2021 – mini-symposium "Mathematical modelling of thermonuclear fusion plasmas"

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Basics on controlled fusion & ITER project

Physical challenges:

- Confinement
- Heat & particle Exhaust
- Disruptions
- Focus on three critical issues:
 - Fast particles: hybrid MHD, turbulence
 - Plasma-wall interaction: the kinetic side
 - Coulomb collisions: towards an efficient (non-linear) operator

Cea Fusion – the principles (I)





Hydrogen isotopes Stable Unstable

- must be produced
- not key for ITER



 $\begin{array}{c|c} \text{Helium} & \text{Neutron} \\ \text{stable } \phi \& \chi & \\ \end{array}$

- impact on device
- used to produce **T**
- not key for ITER



Threshold condition for fusion $E_{th} \ge 10 \text{ keV} \rightarrow Plasma state$ \rightarrow interaction distance 10^{-13} m





Power balance at equilibrium: sources = losses



Ceca Towards ignition: the triple product (Lawson)





Magnetic confinement

T \approx 20 keV(threshold reaction)n \approx 10²⁰m⁻³($\beta = P_{kin} / P_{magn} < 1$)

 $\Rightarrow \text{ Pressure } \sim \text{few atm} \\ (P_{Sun} \sim 10^{11} \text{atm})$

 \Rightarrow Requires $\tau_E \approx 3s$

Cea Particle confinement



Particles confined by strong magnetic field B







Strong anisotropy of the transport: // to B >> ⊥ to B ⇒ Fast homogenization of magnetic flux surfaces

- Magnetic surfaces are (at leading order) isobar, iso-temperature, isocurrent
- Major radius R
- Lorentz force balances
 pressure gradient
 - $\mathbf{j} \times \mathbf{B} = \boldsymbol{\nabla} p$

Nested magnetic surfaces

Cea MHD confinement



Strong anisotropy of the transp
 ⇒ Fast homogenization of magi

 Magnetic surfaces are (at leading order) isobar, iso-temperature, isocurrent

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 - $\mathbf{j}\times\mathbf{B}=\boldsymbol{\nabla}p$



Cea Iter



Milestones ITER Agreement 2006 2007-2009 Land clearing & levelling 2010-2014 Ground structure & seismic foundations 2010-2021 Construction of ITER buildings **Construction of** 2014-2021 **Tokamak Building** Assembly phase 1 2018-2025 Dec 2025 First Plasma First D-T plasma 203X

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- Interface with plasma facing components: in general, hot & diluted plasma
 - \Rightarrow emission of **impurities**
 - \Rightarrow inefficient **pumping**
- Preferable solution: dense, cold
 & radiating plasma
 - → appropriate device = **divertor**
 - Heat exhaust (~10 MW.m⁻²)
 - Pumping ashes (He)
 - Controlling impurity influx
 - Fueling the plasma with D-T
 - Resilience to transient heat loads



$\textcircled{COM} Plasma-wall interaction \rightarrow the divertor$



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Current



- Magneto-hydrodynamic (MHD) instabilities driven by
 - Pressure
 - Current
 - Radiative losses
- Dynamics of MHD instabilities
 - Steady-state
 - Oscillations
 - Disruption







- Pressure
- Current
- Radiative losses
- Dynamics of MHD instabilities
 - Steady-state
 - Oscillations
 - Disruption



Normalized density

Cea Controlling runaway electrons

Disruptions = sudden (few ms) loss of confinement

- Huge forces on plasma vessel
- Runaway Electron beam v_e ~ c \rightarrow up to 100 MJ in ITER
- Strategies: detection, avoidance and/or mitigation

Mitigation of RE at JET with shattered D2 pellets [Reux et al., PRL 2021]





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Tearing MHD mode ↓ Stochastization of field lines ↓ RE no longer confined in 100μs



- Several transport channels
 - Heat $(\tau_E) \rightarrow$ electrostatic turbulence
 - α particles, runaway electrons \rightarrow electromagnetic modes
 - Impurities \rightarrow turbulence & collisions
- Key physics: Resonant and/or non-resonant coupling between particles and EM fields



- Model description: fluid moments and/or adiabatic vs. (gyro)kinetic → coupling issues
- Mix of slow & fast particles: thermal particles vs. α or runaway electrons, ions vs. electrons \rightarrow resolution issues







Cea Kinetic vs. fluid & MHD



(Gyro)Kinetic:

- Vlasov-like equation
- Sources & sinks
- Collisions

$$\mathrm{d}_t F_s = S + C(F_s)$$

- **(5)**6D + time: $F_s(\mathbf{x}, \mathbf{v}, t)$
- Integrals on velocity to get charge & current densities

Fluid & MHD:

 Conservation equations of particles, momentum, energy

$$\rho d_t \mathbf{V} = -\nabla P + \mathbf{J} \times \mathbf{B} + \nu \nabla^2 \mathbf{V}$$
$$\partial_t \mathbf{B} = \nabla \times (\mathbf{V} \times \mathbf{B}) + \frac{\eta}{\mu_0} \nabla^2 \mathbf{B}$$
$$d_t = \partial_t + \mathbf{V} \cdot \nabla$$

+ eq on pressure P

3D + time: V(**x**,t), P(**x**,t), ...



Both descriptions require High Performance Computing

E.g. GK ion turbulence in ITER: a/ ρ_{ci} ~10³ & $\omega_{turb}\tau_{E}$ ~10⁵ \rightarrow ~3.10¹¹ grid points, ~20 Mh.core





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Cea MHD valid at high frequency & large scale



XTOR-K

 $t = 5 \ 150 \ \tau_{\wedge}$

[Lütjens 2010; Nicolas 2012; Brochard 2020]

Fishbone [Brochard 2020]

 $t = 2600 \tau_{\Lambda}$

JOREK



• Works well for ideal global modes $(\omega \tau_A \sim 1)$ $\tau_A = a/v_A \sim \beta^{1/2} a/v_{th}$

[Huysmans 2007; Czarny 2008; Hoelzl 2021] t' = 1.00 ms t' = 1.05 ms

Vertical instability [Artola 2021]

Numerical bottlenecks:

- Implicit schemes to filter out fast time scales
- Small resistivity $(\eta_{\text{ITER}} \sim 10^{-9}) \rightarrow \text{small scales} (\sim \eta^{1/2})$
- Nonlinear dynamics is important (e.g. crashes, ...)

Requires many modes

Inversion of (very) large sparse matrices

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• Accounting for departures from MHD flow is essential for quantitative predictions \rightarrow lon flow $\mathbf{V}_i = \mathbf{V} + \mathbf{V}_* + ...$ MHD flow $\mathbf{V} = \frac{\mathbf{E} \times \mathbf{B}}{B^2} + V_{\parallel}\mathbf{b}$ Diamagnetic flow $\mathbf{V}_* = \frac{\mathbf{B} \times \nabla P_i}{enB^2}$ $mnd_t \mathbf{V}_i = -\nabla P_i + en(\mathbf{E} + \mathbf{V}_i \times \mathbf{B}) - \nabla .\Pi_{gv}$

"Gyroviscous cancellation": gyroviscous stress tensor Π_{gv} Hazeltine,'85] cancels out with diamagnetic dynamic

$$mn\{\partial_t \mathbf{V}_* + [(\mathbf{V} + \mathbf{V}_*) \cdot \boldsymbol{\nabla}]\mathbf{V}_* + \mathbf{b}(\mathbf{V}_* \cdot \boldsymbol{\nabla})V_{\parallel}\} = -\boldsymbol{\nabla} \cdot \boldsymbol{\Pi}_{gv} - \boldsymbol{\nabla}\chi$$

Residual pressure $\chi = (P_i/2\omega_c) \mathbf{b} \cdot \nabla \times \mathbf{V}_{\perp i}$ (~ $\rho_* \mathsf{P}_i <<\mathsf{P}_i$) incorprated in P_{total}

Modified density, momentum, ... eqs $d_t \rho + \rho \nabla \cdot \mathbf{V} + \nabla \cdot (\rho \mathbf{V}_*) = 0$ $\rho d_t \mathbf{V} + \rho (\mathbf{V}_* \cdot \nabla) \mathbf{V}_{\perp} = -\nabla P + \mathbf{J} \times \mathbf{B}$ OK in slab geometry More complex in toroidal [Ramos,'05]

$\underbrace{Ce2}$ Limitations of fluid/MHD \rightarrow Hybrid codes



Fast particles interact resonantly with (**E**,**B**) fields:

modify MHD modes drive Energetic Particle driven modes

- \Rightarrow Track wave-particle resonances
- ⇒ kinetic description
- ▶ Hybrid MHD codes:
 Thermal particles → MHD
 Fast particles → (gyro)kinetic
 ⇒ Coupling via stress tensor



$$\begin{split} \Pi_{hot} &= P_{\parallel} \mathbf{b} \mathbf{b} + P_{\perp} \left(\mathbf{I} - \mathbf{b} \mathbf{b} \right) \\ P_{\parallel} &= \int d^{3} \mathbf{v} F_{h} m v_{\parallel}^{2} \\ P_{\perp} &= \int d^{3} \mathbf{v} F_{h} \frac{1}{2} m v_{\perp}^{2} \end{split}$$
 [Park '92; Briguglio '98; Lütjens '15]





Cea Challenging electron-ion turbulence simu.





Cea Confinement: core-edge coupling





- Edge transport barrier = spontaneous reduced turbulent transport
- Critical role of sheared electric field $v_{E\theta} = -E_r/B \rightarrow \text{vortex stretching}$ (already in L-mode)



- Key challenges when addressing core & edge plasma transport:
 - Closed vs. open field lines ⇒ loss of periodicity, plasma-wall interaction
 - Complex 3D geometry
 - Huge variations of temperature (10 keV → a few eV)



Ceal Physics of plasma-wall interaction = // BC





PWI using immersed boundary conditions 20





SOLEDGE3X-EIRENE



[Caschera 2018; Dif-Pradalier 2020]

GYSELA

Weak collisions... Yet essential





- Critical issues governed by collisions (NS vs. Euler):
 - Relaxation towards Maxwellian
 - Flow damping: friction on trapped particles (→ transition between turbulent regimes, ...), Zonal Flow damping (→ turbulence saturation)
 - Impurity transport, especially high Z + possible synergy turb./coll.
 - Momentum & Energy exchanges between species

General form of collision operator



 Landau formulation (1936) of the Fokker-Planck operator (limit of small angle scattering events & cut-off at Debye shielding distances)

$$\frac{\partial f_a}{\partial t} = \sum_b C_{ab}(f_a, f_b) \qquad \text{Coulomb logarithm: } \ln \Lambda = \ln\left(\frac{\lambda_D}{b_0}\right) \gg 1$$
$$C_{ab}(f_a, f_b) = \frac{\partial}{\partial \mathbf{v}} \cdot \left\{ \frac{2\pi e_a^2 e_b^2 \ln \Lambda}{m_a} \int \mathrm{d}^3 \mathbf{v}' f_a(\mathbf{v}) f_b(\mathbf{v}') \right.$$
$$\frac{\mathbb{I}u^2 - \mathbf{u} \otimes \mathbf{u}}{u^3} \cdot \left(\frac{1}{m_a} \frac{\partial \ln f_a(\mathbf{v})}{\partial \mathbf{v}} - \frac{1}{m_b} \frac{\partial \ln f_b(\mathbf{v}')}{\partial \mathbf{v}'} \right) \right\}$$
$$\mathbf{u} = \mathbf{v} - \mathbf{v}'$$

- Particle, momentum & energy conservation → exchanges between species
- Satisfies Boltzmann H-theorem (1872): Entropy production & C(f_{Maxwell})=0

[Helander-Sigmar'05, Hirschman-Sigmar '76 '77 '81, Hinton-Hazeltine '76, Hinton '83, Rosenbluth-Hazeltine-Hinton '72, Abel '08, Barnes '09, Donnel '19]

Cea Towards a nonlinear collision operator



- ► Usually linearized: $f_s = f_{M,s} + \delta f_s$ with $\delta f_s \ll f_{M,s}$ $C_{ab}(f_a, f_b) \approx C_{ab}(f_{Ma}, f_{Mb}) + C_{ab}(\delta f_a, f_{Mb}) + C_{ab}(f_{Ma}, \delta f_b)$ Test particle op. Field op.
- Full operator required at the edge (strong fluctuations, departure F_{Maxwell})

$$\mathcal{C}_{ab}(f_a, f_b) = -\frac{4\pi e_a^2 e_b^2 \ln \Lambda}{m_a m_b} \frac{\partial}{\partial \mathbf{v}} \cdot \left(\frac{m_a}{m_b} \frac{\partial h_b}{\partial \mathbf{v}} f_a - \frac{1}{2} \frac{\partial^2 g_b}{\partial \mathbf{v} \partial \mathbf{v}} \cdot \frac{\partial f_a}{\partial \mathbf{v}}\right)$$

Physical issues:

 $\nabla_v^2 h_b = -4\pi f_b \qquad \nabla_v^2 g_b = 2h_b$

Rosenbluth potentials (1957)

• Neoclassical results $df_a/dt = C_{ab} \rightarrow \chi_{neo}$, $V_{\theta,neo}$, $\Gamma_{Z,neo}$, ...

Numerical issues:

Conservation properties

- Whatever mass ratio m_a/m_b and charge states e_s
- To be adapted to actual choice of variables, e.g. $(v, v_{\parallel}/v), (v_{\parallel}, \mu)$
- Moderate CPU time → parallelization

Cea Conclusions



- First principal simulations are crucial for ITER (pulse preparation, post-pulse analysis)
- At the limits of current HPC supercomputers (memory & CPU time)
- A hierarchy of complementary approaches

Reduced 1D models3D Fluid, MHD5D Gyrokinetic6D kinetic(D,V), QuasiLinear, K- ε , ...Eulerian, PIC/MC, SL

- ▶ Numerical bottlenecks: coupling \neq approaches, core \rightarrow wall
- Validity limitation of physical models \Rightarrow numerical issues
- Other issues: optimization of 3D magnetic configurations (stellarators), free boundary magnetic equil., ...

Inputs from Maths is essential!

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Back-up slides

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Cea Controlling relaxation oscillations







- ► Pressure / current exceeds instability threshold → fast relaxation → recovery
- Bursts of particle & heat fluxes
- Control is mandatory: via external coils or localized heating / current

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